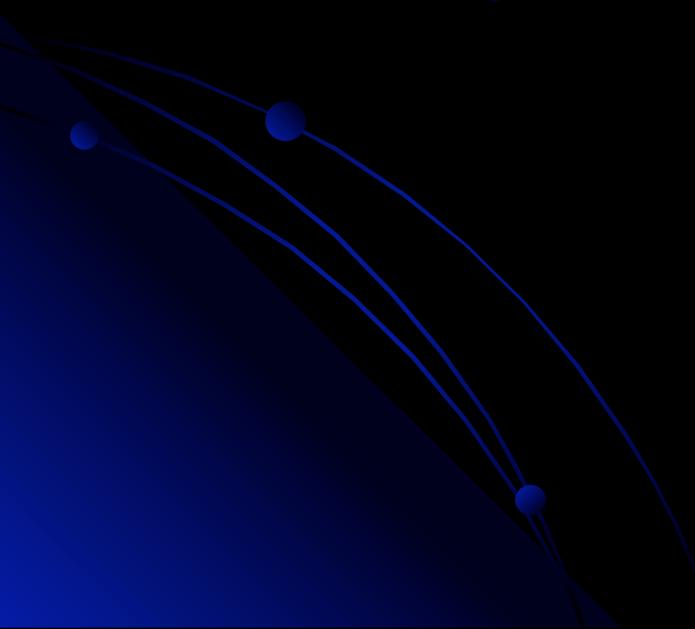


Pre-Launch Testing of NGSLR Ranging to LRO

Anthony Mallama
Jan McGarry
Tom Zagwodzki
Jack Cheek
Christopher Clarke
All at NASA/GSFC

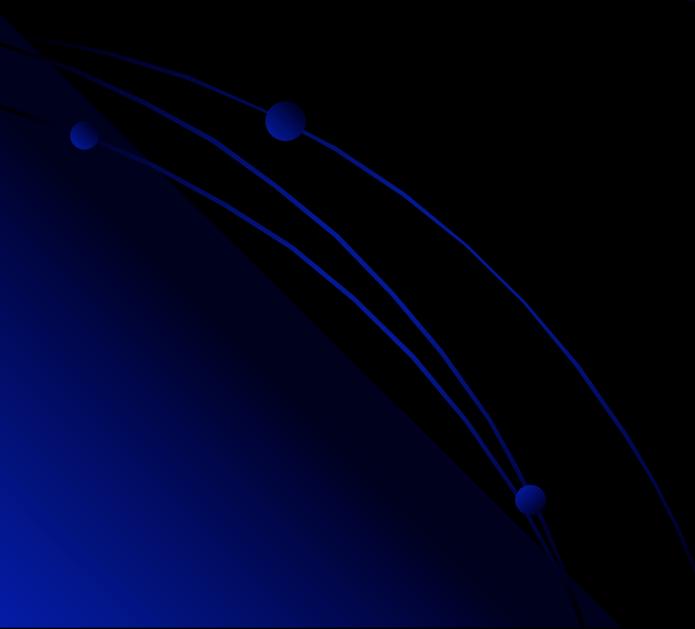
Goal of LRO Laser Ranging

- Generate a precise LRO orbit
- Derive an improved lunar gravity field
- LRO orbit will be the reference for data from the Lunar Orbiter Laser Altimeter (LOLA) instrument
- LOLA provides high accuracy surface mapping



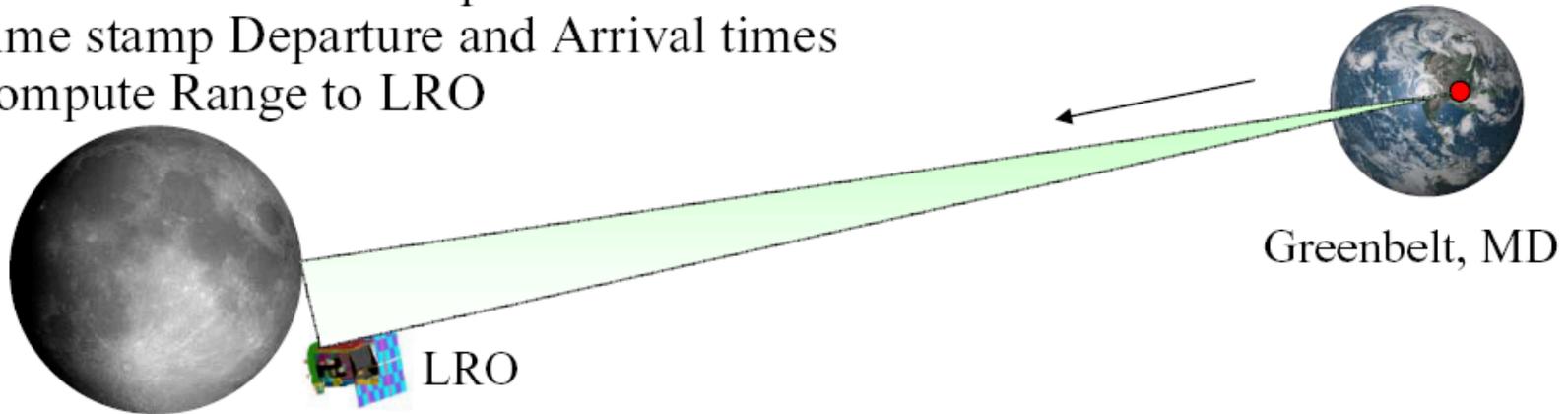
NGSLR to LRO Laser Ranging

- Laser is 28 Hz, 50 milli-Joule, 532 nm
- Laser Ranging receiver telescope on High Gain Antenna
- Fiber optic bundle carries photons to LOLA receiver
- LOLA 28 Hz duty cycle has separate time windows for Earth and Moon pulses



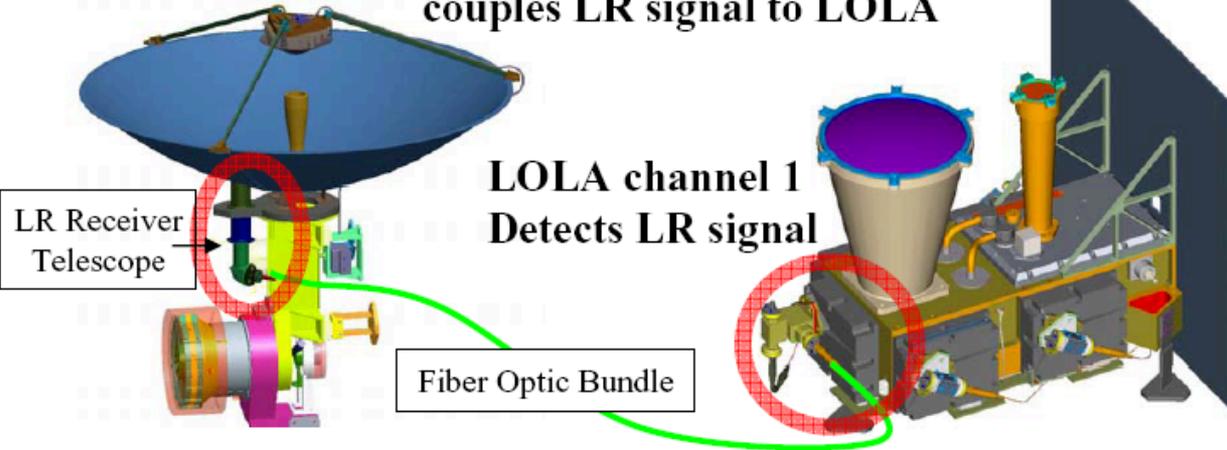
NGSLR to LRO Laser Ranging

- Transmit 532 nm laser pulses at 28 Hz to LRO
- Time stamp Departure and Arrival times
- Compute Range to LRO

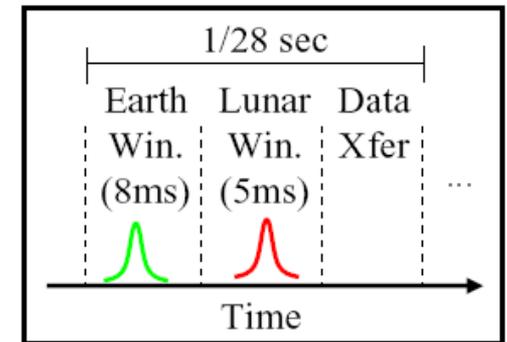


Receiver telescope on HGAS
couples LR signal to LOLA

LOLA channel 1
Detects LR signal



LR Timeshares LOLA Detector
With Lunar surface returns



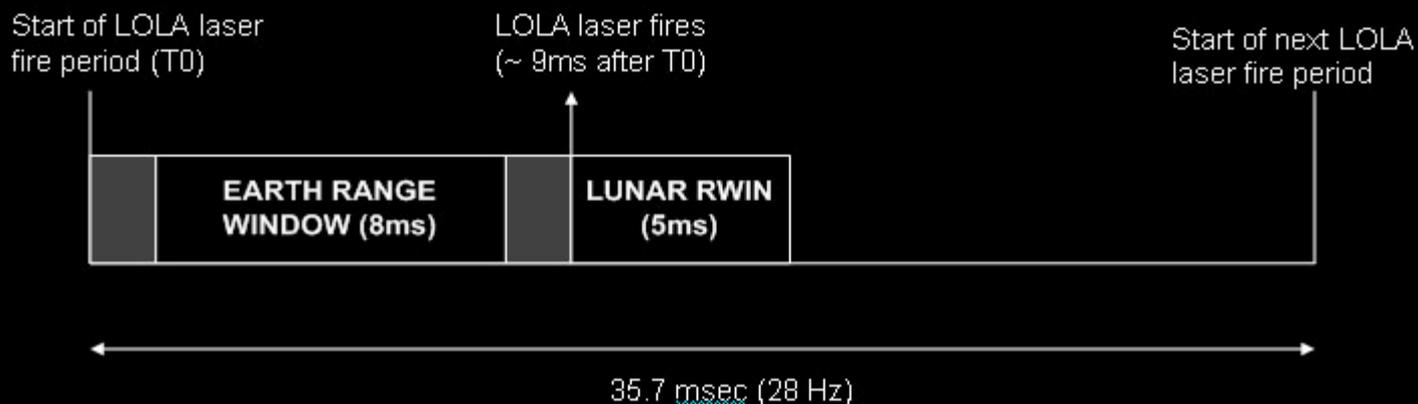
Testing Overview

- NGSLR testing program has 137 elements
- 19 are specific to LRO
- There are 4 general categories of LRO tests:
 - 1) Telescope must accurately point to LRO
 - 2) LRO is scheduled as highest priority target
 - 3) NGSLR software is correctly modified for LRO
 - 4) Laser fire is correctly controlled and measured

Tests Summarized in this Presentation

- Laser pulses arrive at the LR receiver during the LOLA Earth window
 - Laser fire offset and frequency can be manually controlled
 - Telescope pointing is correct and accurate
- 

Laser pulses arrive at the LR receiver during the LOLA Earth window



Laser pulses arrive at the LR receiver during the LOLA Earth window

Considerations for hitting the Earth window

- Modulo of time on spacecraft clock with $1 / 28$ second must be 5 ± 1 milliseconds
- SCLK file relates terrestrial time and spacecraft time
- MET = spacecraft time - offset
- Light-time of NGSLR-to-LRO range

Laser pulses arrive at the LR receiver during the LOLA Earth window

SCLK File:

\begindata

SCLK_KERNEL_ID = (@2008-06-27/20:37:49.00)

SCLK_DATA_TYPE_85 = (1)

SCLK01_TIME_SYSTEM_85 = (2)

SCLK01_N_FIELDS_85 = (2)

SCLK01_MODULI_85 = (4294967296 65536)

SCLK01_OFFSETS_85 = (0 0)

SCLK01_OUTPUT_DELIM_85 = (1)

SCLK_PARTITION_START_85 = (0.000000000000000E+00)

SCLK_PARTITION_END_85 = (2.8147497671065E+14)

SCLK01_COEFFICIENTS_85 = (

0.000000000000000E+00

3.1579264184000E+07

1.0000000010000E+00

1.5485503275008E+13

2.6786929218400E+08

1.0000000010000E+00)

\begintext



Ticks



Seconds



Rate

Laser pulses arrive at the LR receiver during the LOLA Earth window

Determine the fractional part of MET seconds:

Convert UT to ET

Find the interval for the ET in the SCLK file

(ET = 0 corresponds to epoch of J2000)

$\Delta ET = ET - \text{Seconds at start of interval}$

$\Delta \text{Ticks} = \Delta ET / \text{Rate at start of interval}$

Add ΔTicks to Ticks at the start of interval

Spacecraft seconds = $\text{Ticks} / 65536$ (nominal rate)

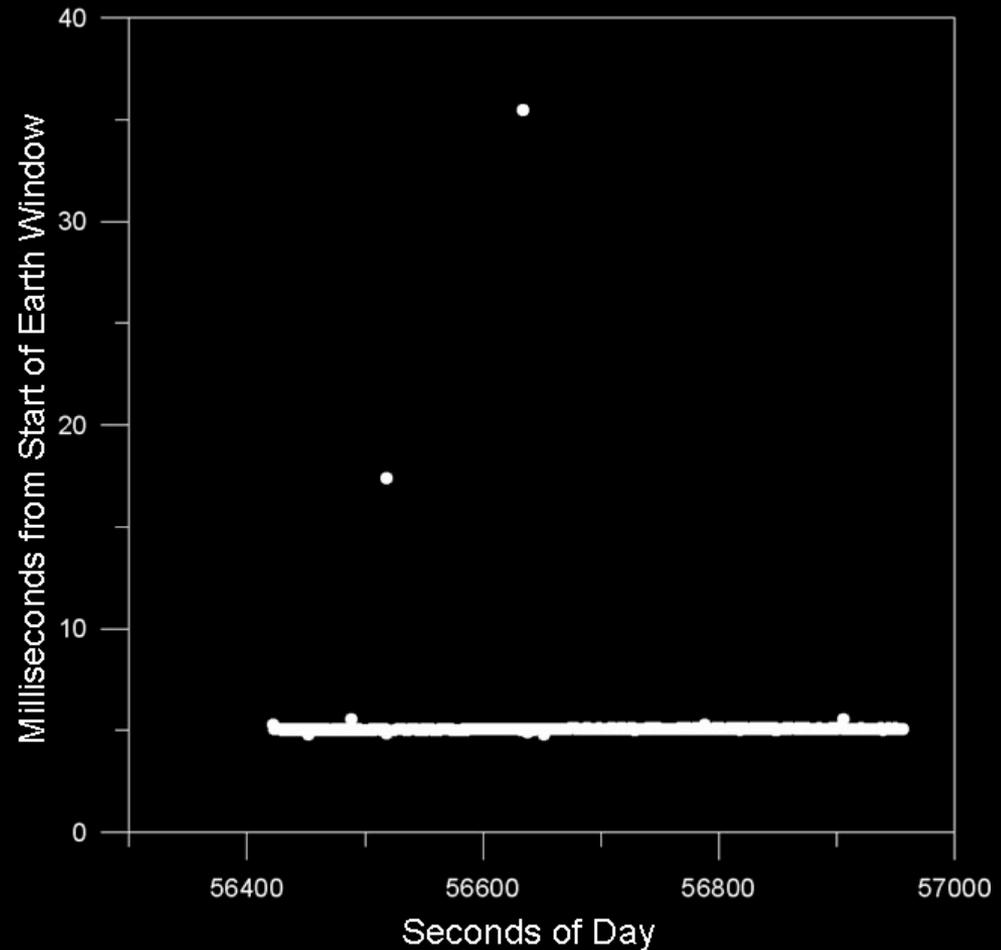
Subtract the Offset

Add the NGSLR-to-LRO range

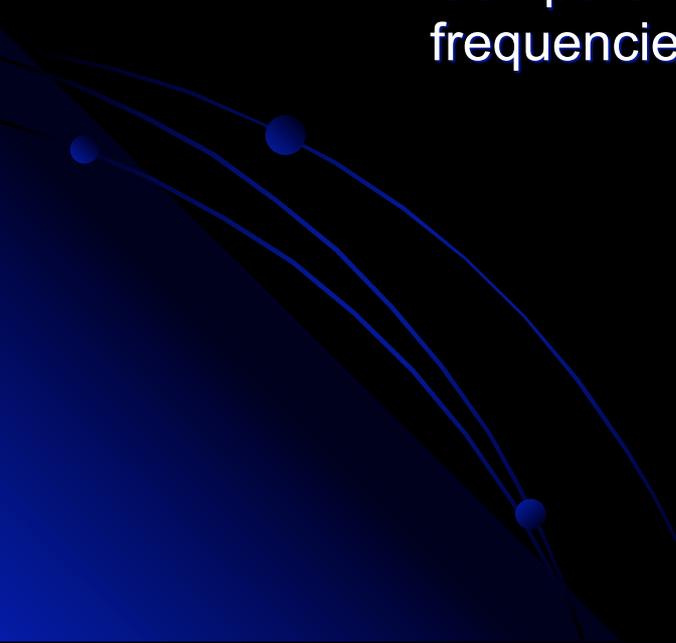
Take the modulus with $1 / 28$ second

Laser pulses arrive at the LR receiver during the LOLA Earth window

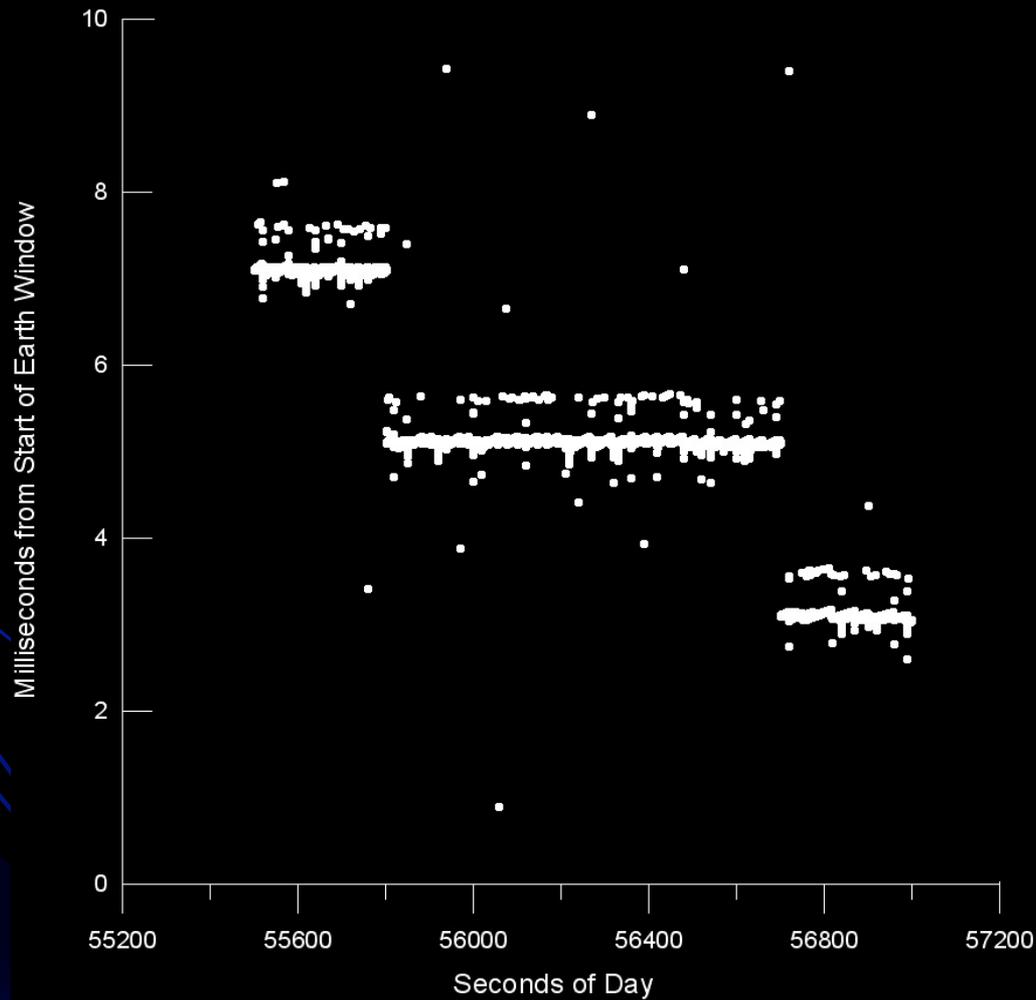
99.9% are within
5 +/- 1 msec.



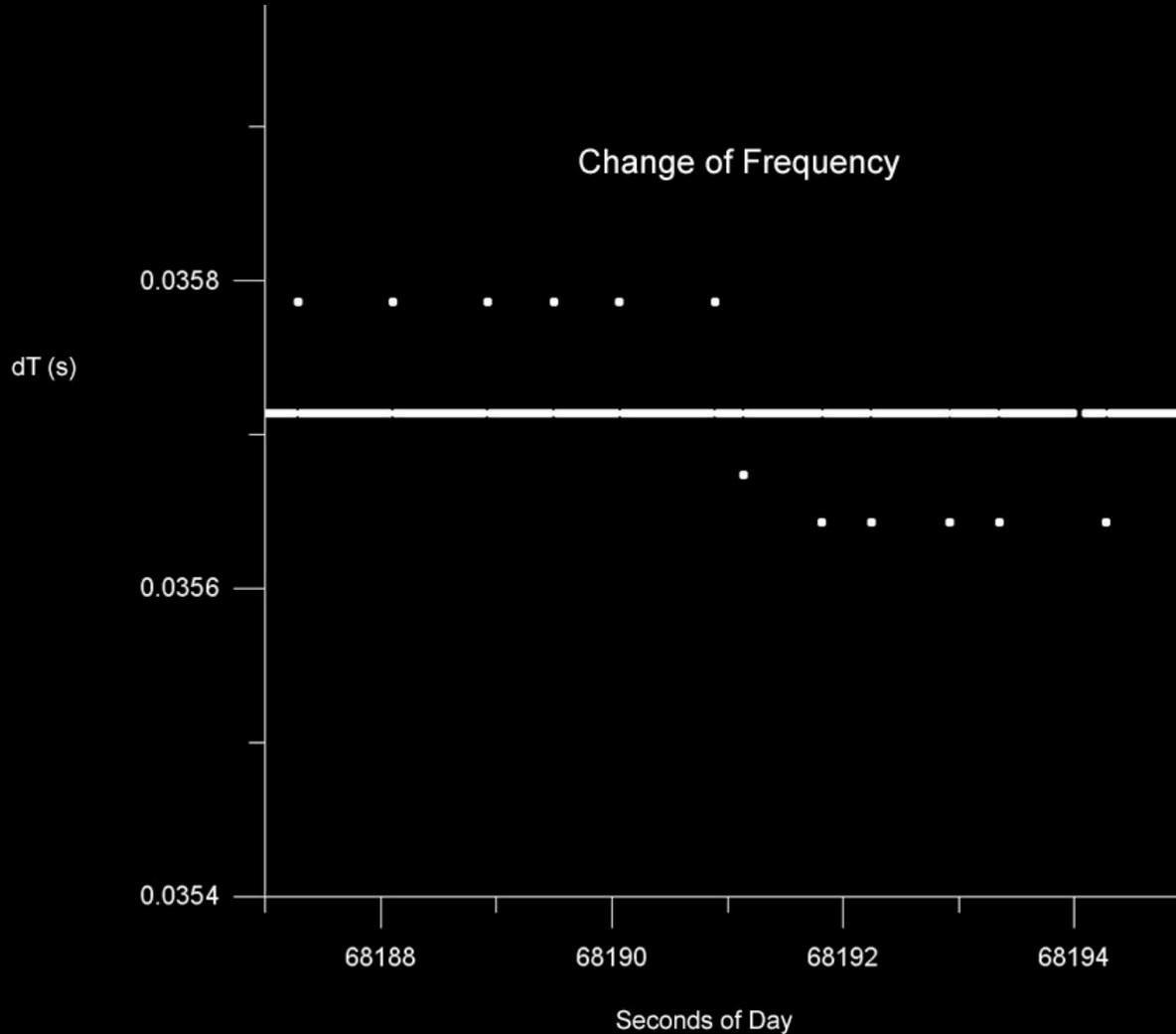
Laser fire offset and frequency can be manually controlled

- Command offsets from -35.7 ms to +35.7 ms
 - Command frequency changes of +/- 100 us/s
 - Analyze commanded recorded fire times
 - Compare commanded and measured offsets and frequencies
- 

Laser fire offset and frequency can be manually controlled



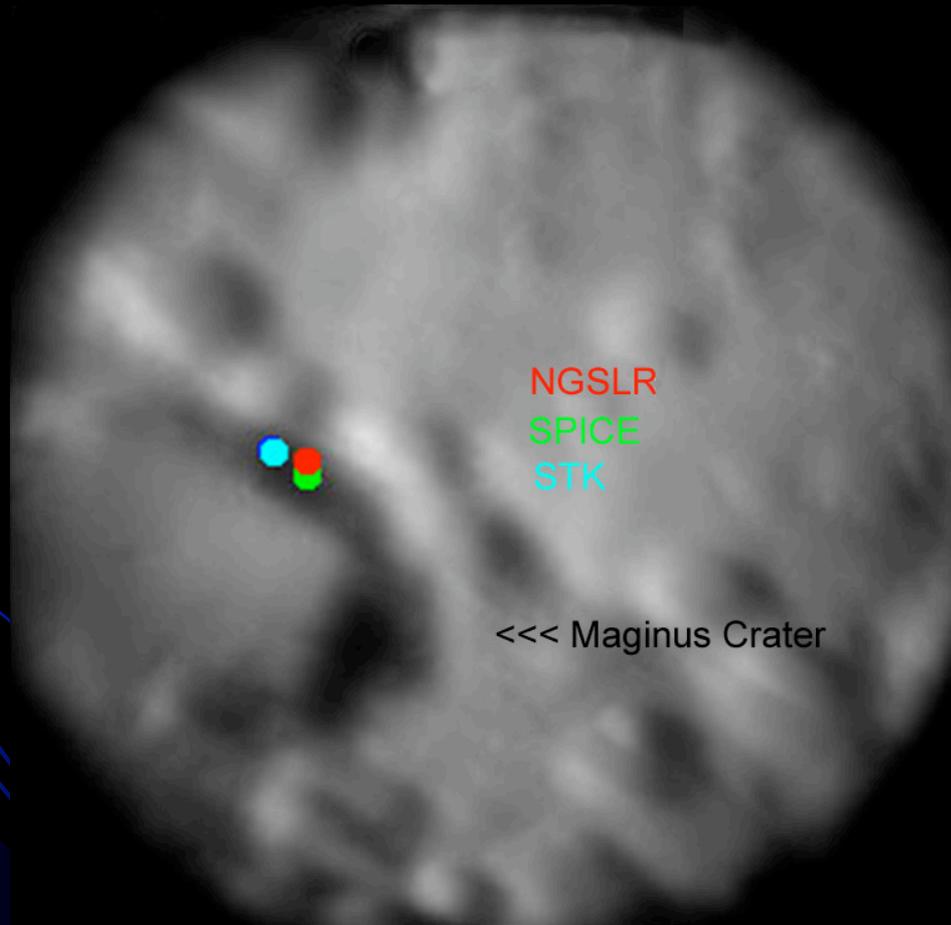
Laser fire offset and frequency can be manually controlled



Telescope pointing is correct and accurate

- Simulated an LRO pass
- Used NASA Flight Dynamics Facility predictions
- Acquired images of Moon with NGSLR camera
- Marked telescope pointing on the NGSLR images
- Compared telescope pointing with:
 - STK image data provided by FDF
 - Lunar coordinates generated by a SPICE program

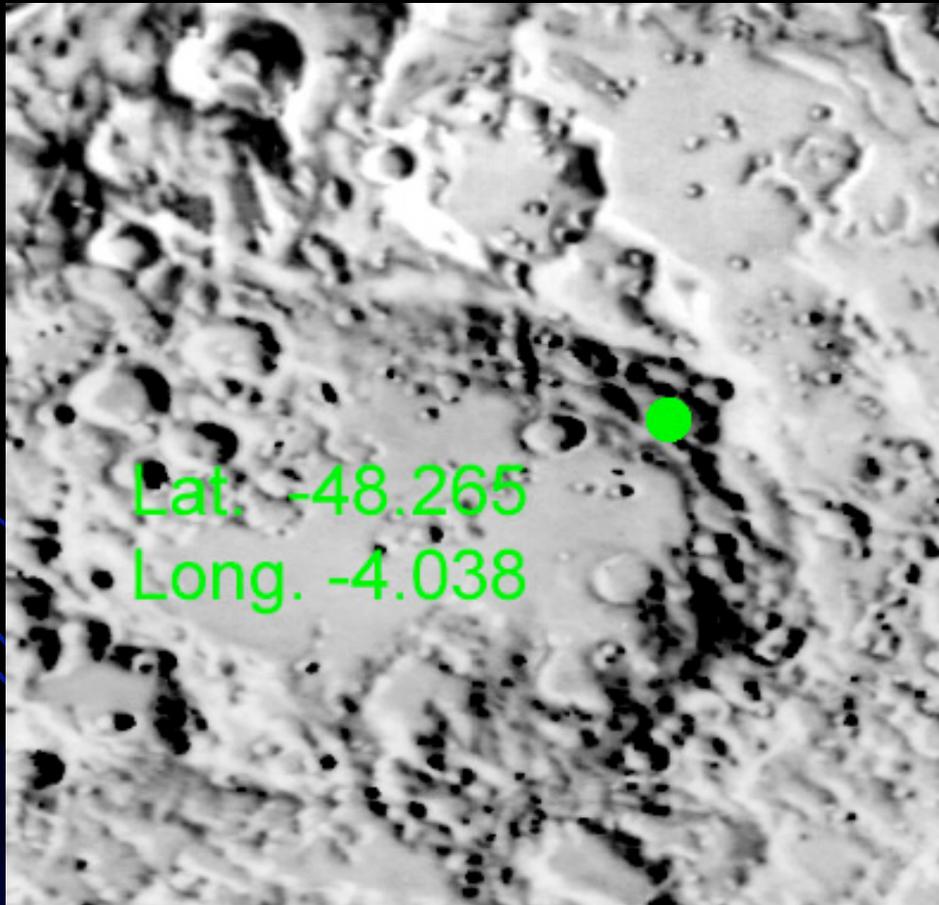
Telescope pointing is correct
and accurate



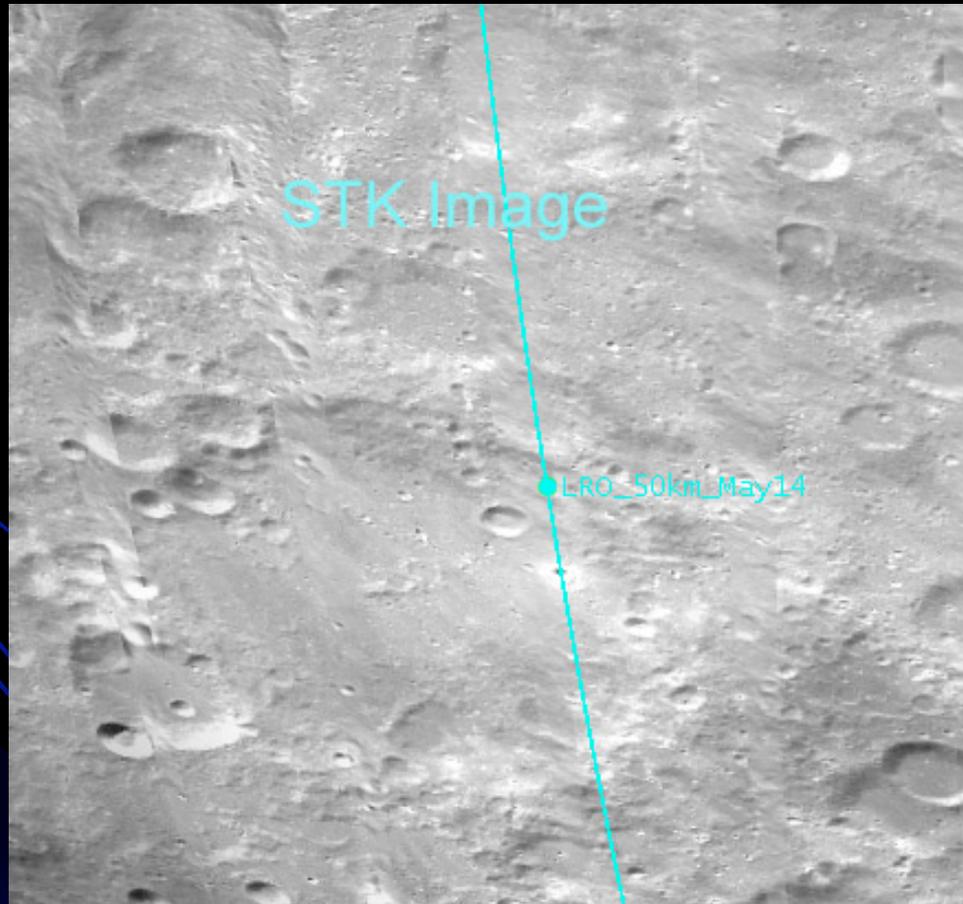
Telescope pointing is correct and accurate

- SPICE Kernels:
 - NGSLR position, custom generated kernel
 - Earth and Moon positions, DE421
 - LRO position, special orbit kernel from FDF
 - Lunar 'Mean Earth' reference frame
- Point-ahead correction (SPICE: XCN+S)
 - Transmission
 - Converged Newtonian light time correction
 - Stellar aberration
- Plotted lunar coordinates
 - USGS airbrushed shaded relief map
 - Warped to ULCN2005

Telescope pointing is correct
and accurate

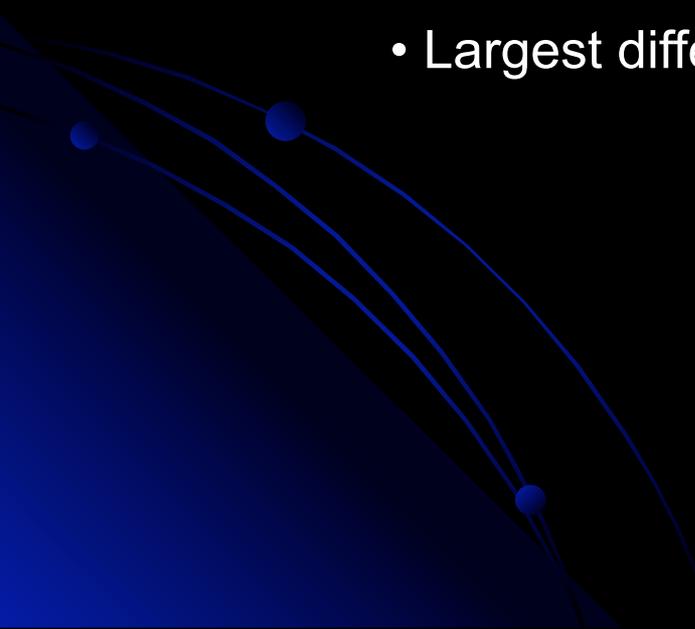


Telescope pointing is correct and accurate



Telescope pointing is correct and accurate

- NGSLR versus SPICE \rightarrow 2.4 arc s (rms)
- Largest difference \sim 4 arc s
- NGSLR versus STK \rightarrow 3.2 arc s (rms)
- Largest difference \sim 5 arc s



Summary

- The LRO portion of the NGSLR testing program is nearly complete
 - The telescope will track LRO correctly and accurately
 - LRO will be scheduled as the highest priority target
 - NGSLR software is correctly modified for LRO
 - Laser fire is correctly controlled and measured
- 